JAMES WEBB SPACE TELESCOPE

GODDARD SPACE FLIGHT CENTER

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The most popular questions about Webb targeted at the general interest level.

Frequently Asked Questions Lite

What is the James Webb Space Telescope?

The James Webb Space Telescope, also called Webb or JWST, is a large, space-based observatory, optimized for infrared wavelengths, which will complement and extend the discoveries of the Hubble Space Telescope. It launches in 2019

It will cover longer wavelengths of light than Hubble and will have greatly improved sensitivity. The longer wavelengths enable JWST to look further back in time to see the first galaxies that formed in the early universe, and to peer inside dust clouds where stars and planetary systems are forming today.

More Info

What are the most exciting things we will learn?

We have yet to observe the era of our universe's history when galaxies began to form. We have a lot to learn about how galaxies got supermassive black holes in their centers, and we don't really know whether the black holes caused the galaxies to form or vice versa. We can't see inside dust clouds with high resolution, where stars and planets are being born nearby, but Webb will be able to do just that. We don't know how many planetary systems might be hospitable to life, but Webb could tell whether some Earth-like planets have enough water to have oceans. We don't know much about dark matter or dark energy, but we are expecting to learn more about where the dark matter is now, and we hope to learn the history of the acceleration of the universe that we attribute to dark energy. And then, there are the surprises we can't imagine!

More Info

Why is Webb an infrared telescope?

By viewing the universe at infrared wavelengths Webb will show us things never before seen by any other telescope. It is only at infrared wavelengths that we can see the first stars and galaxies forming after the Big Bang. And it is with infrared light that we can see stars and planetary systems forming

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FAQS:

Vital Facts

Fact sheet on the Webb Mission.

FAQ Lite

The most popular questions about Webb. (General Public)

FAQ Full

All the major aspects of the Webb Mission are covered here. (General Public)

Technical FAQ

Technical FAQ on a variety of mission issues, aspects and capabilities. (Science/Technical)

Solar System Observations FAQ

Technical FAQ specifically on Solar System observations. (Science/Technical)

Q&A:

Q & A with engineers cryo-testing the telescope. (new window)

Lee Feinberg AMA: Optical Telescope Element Manager for Webb (new window)

Lee Feinberg talks about the top 3 things that Webb can do that Hubble cannot and more.

Paul Geithner: Freezing NASA's Webb Telescope Is a Matter of Survival

(new window)

Paul Geithner (Webb Deputy Project Manager, Technical) provides answers to questions about the kind of freezing temperatures the Webb telescope will endure in space. (General Public) inside clouds of dust that are opaque to visible light.

The primary goals of Webb are to study galaxy, star and planet formation in the universe. To see the very first stars and galaxies that formed in the early universe, we have to look deep into space to look back in time (because it takes light time to travel from there to here, the farther out we look, the further we look back in time).

The universe is expanding, and therefore the farther we look, the faster objects are moving away from us, redshifting the light. Redshift means that light that is emitted as ultraviolet or visible light is shifted more and more to redder wavelengths, into the near- and mid-infrared part of the electromagnetic spectrum for very high redshifts. Therefore, to study the earliest star and galaxy formation in the universe, we have to observe infrared light and use a telescope and instruments optimized for this light.

Star and planet formation in the local universe takes place in the centers of dense, dusty clouds, obscured from our eyes at normal visible wavelengths. Near-infrared light, with its longer wavelength, is less hindered by the small dust particles, allowing near-infrared light to escape from the dust clouds. By observing the emitted near-infrared light we can penetrate the dust and see the processes leading to star and planet formation.

Objects of about Earth's temperature emit most of their radiation at mid-infrared wavelengths. These temperatures are also found in dusty regions forming stars and planets, so with mid-infrared radiation we can see the glow of the star and planet formation taking place. An infrared-optimized telescope allows us to penetrate dust clouds to see the birthplaces of stars and planets.

More Info

Will Webb take amazing pictures like Hubble? Can Webb see visible light?

YES! We are going to be looking at things we've never seen before and looking at things we have seen before in all new ways.

The beauty and quality of an astronomical image depends on two things: the sharpness and the number of pixels in the camera. On both of these counts, Webb is very similar to, and in many ways better than, Hubble. Although Webb images will be infrared, this can be translated by computer into a visible picture (just like we have done with Spitzer, which has produced beautiful pictures as well). Additionally Webb can see orange and red visible light. Webb images will be different, but just as beautiful as Hubble's. Read More Also listen to this interview with John Mather about "Making Data Beautiful" on the NASA Blueshift podcast.

What will Webb's first targets be?

Paul Geithner: Components & Structure of Webb

(new window)

The Webb telescope is a showcase for new technologies. Recently Paul Geithner provided a closer look at the technologies on the observatory. (General Public)

Paul Geithner: It's All About Infrared - Why Build Webb (new window)

Paul Geithner provides insight on why the Webb telescope focuses on the infrared. (General Public)

Paul Geithner: James Webb Space Telescope Coming Together (new window)

Paul Geithner discusses progress, plans and next steps in building the Webb Telescope. (General Public)

John Mather: "Ask Me Anything"

Dr. John Mather (Nobel Laureate and Webb Senior Project Scientist) answering questions on Reddit. (General Public)

John Mather: Webb Q&A 1

Dr. John Mather captured on Twitter during our first Tweet Chat. (General Public)

John Mather: Webb Q&A 2

Dr. John Mather captured on Twitter during our second Tweet Chat. (General Public)

Mark Clampin: Exoplanets

Dr. Mark Clampin (Webb Observatory Project Scientist) answers questions about Webb and exoplanets.

SXSW: Finding Life on Other Planets

During SXSW 2014, we held a tweet-chat with some of the scientists on the "First Signs: Finding Life on Other Planets" panel.

Women of Webb

Women with diverse jobs on the James Webb Space Telescope answered questions about the female experience working on a NASA flagship mission in the TwoXChromosomes subreddit.

Focus on XSTEM

During the USA Science & Engineering Festival in April 2014 and focused on STEM (Science, Technology, Engineering, Math) topics.

The first targets for Webb will be determined through a process similar to that used for the Hubble Space Telescope and will involve NASA, ESA, CSA and scientific community participants.

The first engineering target will come before the first science target and will be used to align the mirror segments and focus the telescope. That will probably be a relatively bright star or possibly a star field.

Sara Seeger: Search for Life Part 1 | Part 2

MIT's Dr. Sara Seager answered questions about exoplanets, the search for life, and the next technologies (like Webb!)

How does Webb compare with Kepler?

The **Kepler mission** is designed to answer a simple question. What fraction of stars have terrestrial planets located in or near the habitable zone?

The habitable zone is the region around a star where water can exist on a planet in liquid form. Kepler seeks to answer this question by staring at a small region of the sky containing more than 100,000 stars for 3.5 years or more to look for transiting terrestrial planets, and thus determine what fraction of stars have terrestrial planets. In answering this question, Kepler is generating a large database of confirmed transiting planets together with some of their basic properties. Once we have found these planets, we need the tools to study their physical properties and the composition of their atmospheres. It is Webb that provides the specialized tools to undertake these studies. Kepler is designed to be a "wide and shallow" survey telescope, while Webb is designed for "narrow and deep" focused studies with near and mid-IR imaging and spectroscopy.

More Info

How does Webb compare to the Transiting Exoplanet Survey Satellite (TESS)?

Webb is a very large observatory designed to address a variety of questions across many areas of astrophysics, while TESS concentrates on identifying a large sample of small planets where follow-up observations are feasible with current and planned telescopes. TESS will identify small planets and measure their sizes. Through follow-up observations, we can determine the masses of some of these planets. With both mass and size measurements, we can determine the planets' densities and start to understand what they are made of. This work will provide a foundation for future missions in the search for potentially habitable planets.

TESS will expand on the NASA Kepler mission's census of exoplanets by targeting closer, brighter stars, where follow-up observations are easier to make. The stars TESS studies will on average be 30 to 100 times brighter than the stars earlier Kepler surveyed.

How does Webb compare with Herschel?

Webb and Herschel are complementary.

Webb will be a 6.5m telescope sensitive from gold-colored visible light to the mid-infrared, at wavelengths ranging from 0.6 micrometers to 28.5 micrometers. Herschel was a 3.5m telescope sensitive in the far-infrared from 55 to 670 micrometers wavelength. By working at longer wavelengths, Herschel saw colder objects, such as the earliest stages of star formation in dark clouds and emission from molecules such as water. Webb will view more energetic phenomena including forming proto-stars and very distant galaxies. Getting data with both telescopes on the same objects will build a more complete picture of the astrophysical processes.

Webb is designed to look deeper into space to see the earliest stars and galaxies that formed in the universe and to look deep into nearby dust clouds to study the formation of stars and planets.

In order to do this, Webb will have a much larger primary mirror than Hubble (2.5 times larger in diameter, or about 6 times larger in area), giving it more light-gathering power. It also will have infrared instruments with longer wavelength coverage and greatly improved sensitivity than Hubble. Finally, Webb will operate much farther from Earth, maintaining its extremely cold operating temperature, stable pointing and higher observing efficiency than with the Earth-orbiting Hubble.

More Info

What will Webb tell us about exoplanets? Will it take photos of exoplanets?

Webb will be able to tell us the composition of the atmospheres of exoplanets. It will observe planetary atmospheres through the transit technique. A transit is when a planet moves across the disc of its parent star. Webb will also carry coronographs to enable photography of exoplanets near bright stars (if they are big and bright and far from the star), but they will be only "dots," not grand panoramas. Consider how far away exoplanets are from us, and how small they are by comparison to this distance! We didn't even know what Pluto really looked like until we were able to send an observatory to fly right near it, and Pluto is in our own solar system!

More Info

Will we image objects in our own solar system?

Yes! Webb will be able to observe the planets at or beyond the orbit of Mars, satellites, comets, asteroids, and Kuiper belt objects.

Many important molecules, ices, and minerals have strong characteristic signatures at the wavelengths Webb can observe.

Webb will also monitor the weather of planets and their moons.

Because the telescope and instruments have to be kept cold, Webb's protective sunshield will be blocking the inner solar system from view. This means that the Sun, Earth, Moon, Mercury, and Venus, and of course sun-grazing comets and many known near-Earth objects cannot be observed.

More Info

How far back will Webb see?

Webb will be able to see what the universe looked like around a quarter of a billion years (possibly back to 100 million years) after the Big Bang, when the first stars and galaxies started to form.

More Info

When will Webb launch and how long is the mission duration?

Webb will launch in 2019 from French Guiana on a European Space Agency Ariane 5 rocket. Webb's mission lifetime after launch is designed to be at least 5-1/2 years, and could last longer than 10 years. The lifetime is limited by the amount of fuel used for maintaining the orbit, and by the possibility that Webb's components will degrade over time in the harsh environment of space.

In the first hour: Starting at liftoff, the Ariane rocket will provide thrust for a little over 8 minutes. Webb will separate from the Ariane V launch vehicle a half hour after launch and we will deploy the solar array immediately afterward. We will also release several systems that were locked for launch.

In the first day: Two hours after launch we will deploy the high gain antenna. About ten and a half hours after launch, Webb will pass the Moon's orbit, nearly a quarter of the way to L2. Twelve hours after launch there will be the first trajectory correction maneuver by small rocket engines aboard Webb itself.

In the first week: The second trajectory correction maneuver will take place at 2.5 days after launch. We will start the sequence of major deployment just after that. The first deployments are the fore and aft sunshield pallets, followed by the release of remaining sub-system launch locks. The next deployment is the telescope in which the telescope and the spacecraft bus move apart from each other by about 2 meters when the deployable tower assembly extends. The full sunshield deployment with unfolding and tensioning of the membranes can then be initiated. At 6 days we deploy the secondary mirror, followed by the side wings of the primary mirror.

In the first month: As the telescope cools down in the shade of the deployed sunshield, we will turn on the warm electronics and initialize the flight software. At the end of the first month, we will do the mid-course correction that ensures that Webb will achieve its final orbit around L2. Although the telescope cools to near its operating temperature, the ISIM is warmed with electric heaters to prevent condensation on the instruments as residual water trapped in the materials making up the observatory escapes to the vacuum of space.

In the second month: At 33 days after launch we will turn on and operate the Fine Guidance Sensor, then NIRCam and NIRSpec. The first NIRCam image will be of a crowded star field to make sure that light gets through the telescope into the instruments. Since the primary mirror segments will not yet be aligned, the picture will still be out of focus. At 44 days after launch we will begin the process of adjusting the primary mirror segments, first identifying each mirror segment with its image of a star in the camera. We will also focus the secondary mirror.

In the third month: From 60 to 90 days after launch we will align the primary mirror segments so that they can work together as a single optical surface. We will also turn on and operate the MIRI. By the end of the third month we will be able to take the first science-quality images. Also by this time, Webb will complete its initial orbit around L2.

In the fourth through the sixth month: At about 85 days after launch we will have completed the optimization of the telescope image in the NIRCam. Over the next month and a half we will optimize the image for the other instruments. We will test and calibrate all of the instrument capabilities by observing representative science targets.

After six months: Webb will begin its science mission and start to conduct routine science operations.

More Info

Is Webb serviceable?

Webb will be operated at the second Sun-Earth Lagrange point (L2), located approximately 1 million miles (1.5 million km) away from the Earth, and will therefore be beyond the reach of any manned vehicle currently being planned for the next decade.

Hubble is in low-Earth orbit, located approximately 375 miles (600 km) away from the Earth, and is therefore readily accessible for servicing. In the early days of the Webb project, studies were conducted to evaluate the benefits, practicality and cost of servicing Webb either by human space flight, by robotic missions, or by some combination such as retrieval to low-Earth orbit. Those studies concluded that the potential benefits of servicing do not offset the increases in mission complexity, mass and cost that would be required to make Webb serviceable, or to conduct the servicing mission itself. Hubble is an exception and not the rule. No other satellites but Hubble are serviceable currently.

Is there a danger from micrometeorites?

All of Webb's systems are designed to survive micrometeoroid impacts.

We tested beryllium discs for micrometeoroids using test facilities in the US and showed the micrometeoroids have negligible effects on the beryllium. Cryogenic beryllium mirrors have been flown in space exposed to micrometeoroids without problems. The Spitzer Space Telescope, launched in 2003, has a beryllium primary mirror.

Why is the mirror segmented?

Webb needs to have an unfolding mirror because the mirror is so large that it otherwise cannot fit in the launch shroud of currently available rockets.

The mirror has to be large in order to see the faint light from the first star-forming regions and to see very small details at infrared wavelengths. Designing, building and operating a mirror that unfolds is one of the major technological developments of Webb. Unfolding mirrors will be necessary for future missions requiring even larger mirrors, and will find application in other scientific, civil and military space missions.

More Info

Why are the mirrors hexagonal?

The hexagonal shape allows a segmented mirror to be constructed without gaps that can be roughly circular in shape and needs only 3 variations in prescription.

The hexagonal shape allows a segmented mirror with "high filling factor and six-fold symmetry". High filling factor means the segments fit together without gaps. If we had circular segments, there would be gaps between them. Symmetry is good because we only need 3 different optical prescriptions for all 18 segments, (6 of each prescription). Finally, we want a roughly circular overall mirror shape because that focuses the light into the most compact region on the detectors. An oval mirror, for example, would give images that are elongated in one direction. A square mirror would send a lot of the light out of the central region.

More Info

Why does the sunshield have 5 layers?

Webb has a giant, tennis-court sized sunshield, made of five, very thin layers of an insulating film called Kapton. Why five? One big thick sunshield would conduct the heat from the bottom to the top more than would a shield with five layers separated by vacuum. With five layers to the sunshield, each successive one is cooler than the one below. The heat radiates out from between the layers, and the vacuum between the layers is a very good insulator. From studies done early in the mission development five layers were found to provide sufficient cooling. More layers would provide additional cooling, but would also mean more mass and complexity. We settled on five because it gives us enough cooling with some "margin" or a safety factor, and six or more wouldn't return any additional benefits.

Webb has a large sunshade to protect the telescope from heating by direct sunlight, allowing it to cool down to a temperature below 50 Kelvin (-223° C or -370° F) by passively radiating its heat into space. (Fun fact: You could nearly boil water on the hot side of the sunshield and it's cold enough to actually freeze nitrogen on the cold side.)

Webb is a reflecting telescope that uses 3 curved mirrors. (It's a 3 mirror anastigmat). There is a technical description of this kind of telescope **on Wikipedia**.

Why not assemble it in orbit?

Various scenarios were studied and assembling on-orbit was determined to be unfeasible.

In the very earliest stages of feasibility studies, we examined the possibility of on-orbit assembly for what was then called the next generation space telescope or NGST. At that time, and still today, the International Space Station does not have the capability to assemble precision optical structures. Additionally, the environment around the Space Station is not suitable for the exposed mirrors that Webb has and would have had the possibility to damage or contaminate the optics. The deployment of Webb happens far above Low Earth Orbit and the debris that resides there. Finally, if the Space Station were used as a stopping point for the observatory we would have needed a second rocket to launch it to its final destination at L2. The observatory would have to be designed with much more mass to withstand this "second launch" leaving less mass for the mirrors and science instruments.

Who is James Webb?

This space-based observatory is named after James E. Webb (1906- 1992), NASA's second administrator. Webb is best known for leading Apollo, a series of lunar exploration programs that landed the first humans on the Moon. However, he also initiated a vigorous space science program that was responsible for more than 75 launches during his tenure, including America's first interplanetary explorers. Read more about James E. Webb.







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